# Regularity Criteria of the Solutions to Axisymmetric Magnetohydrodynamic System 

Xuemeng Liu ${ }^{1, \dagger}$


#### Abstract

In this paper, we consider Cauchy problem of the axially symmetric Magnetohydrodynamic (MHD) system. By using energy method, we establish some regularity criteria of the solutions for the axisymmetric solutions of the three dimensional incompressible MHD system.


Keywords Axisymmetric solutions, Regularity criterion, Incompressible magnetohydrodynamics.

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## 1. Introduction

In this paper, we consider Cauchy problem of the incompressible MHD system:

$$
\left\{\begin{array}{l}
\partial_{t} u+(u \cdot \nabla) u-\triangle u+\nabla p=(v \cdot \nabla) B  \tag{1.1}\\
\partial_{t} B+(u \cdot \nabla) B-\triangle B=(B \cdot \nabla) u, \\
\operatorname{div} u=0, \operatorname{div} B=0 \\
(u, B)(\mathbf{x}, 0)=\left(u_{0}, B_{0}\right)
\end{array}\right.
$$

where $u=u(\mathbf{x}, t) \in \mathbb{R}^{3}$ denotes the velocity of the fluid, $B=B(\mathbf{x}, t) \in \mathbb{R}^{3}$ stands for the magnetic field, the scalar function $p=p(\mathbf{x}, t)$ is pressure and $\mathbf{x}=(x, y, z) \in \mathbb{R}^{3}$.

Our main concern here is to establish a family of unique solutions of system (1.1) with the form

$$
\begin{aligned}
u(\mathbf{x}, t) & =u_{r}(r, z, t) e_{r}+u_{\theta}(r, z, t) e_{\theta}+u_{z}(r, z, t) e_{z} \\
B(\mathbf{x}, t) & =B_{r}(r, z, t) e_{r}+B_{\theta}(r, z, t) e_{\theta}+B_{z}(r, z, t) e_{z}
\end{aligned}
$$

in the cylindrical coordinate system. Here,

$$
\begin{aligned}
& e_{r}=\left(\frac{x}{r}, \frac{y}{r}, 0\right), e_{\theta}=\left(-\frac{y}{r}, \frac{x}{r}, 0\right), e_{z}=(0,0,1), \\
& r=\sqrt{x^{2}+y^{2}},(x, y, z)=(r \cos \theta, r \sin \theta, z)
\end{aligned}
$$

[^0]In terms of $\left(u_{r}, u_{\theta}, u_{z}, B_{r}, B_{\theta}, B_{z}\right)$, the MHD system (1.1) can be rewritten as

$$
\left\{\begin{array}{l}
\partial_{t} u_{r}+(u \cdot \nabla) u_{r}-\triangle u_{r}+\frac{u_{r}}{r^{2}}-\frac{u_{\theta}^{2}}{r}+\partial_{r} p-(B \cdot \nabla) B_{r}+\frac{B_{\theta}^{2}}{r}=0,  \tag{1.2}\\
\partial_{t} u_{\theta}+(u \cdot \nabla) u_{\theta}-\triangle u_{\theta}+\frac{u_{\theta}}{r^{2}}+\frac{u_{r} u_{\theta}}{r}-(B \cdot \nabla) B_{\theta}-\frac{B_{r} B_{\theta}}{r}=0, \\
\partial_{t} u_{z}+(u \cdot \nabla) u_{z}-\triangle u_{z}+\partial_{z} p-(B \cdot \nabla) B_{z}=0, \\
\partial_{t} B_{r}+(u \cdot \nabla) B_{r}-\triangle B_{r}+\frac{B_{r}}{r^{2}}-(B \cdot \nabla) u_{r}=0, \\
\partial_{t} B_{\theta}+(u \cdot \nabla) B_{\theta}-\triangle B_{\theta}+\frac{B_{\theta}}{r^{2}}+\frac{u_{\theta} B_{r}}{r}-(B \cdot \nabla) u_{\theta}-\frac{B_{\theta} u_{r}}{r}=0, \\
\partial_{t} B_{z}+(u \cdot \nabla) B_{z}-\triangle B_{z}-(B \cdot \nabla) u_{z}=0 \\
\frac{u_{r}}{r}+\partial_{r} u_{r}+\partial_{z} u_{z}=0 \\
\frac{B_{r}}{r}+\partial_{r} B_{r}+\partial_{z} B_{z}=0 \\
\left(u_{r}, u_{\theta}, u_{z}, B_{r}, B_{\theta}, B_{z}\right)(r, z, 0)=\left(u_{0}^{r}, u_{0}^{\theta}, u_{0}^{z}, B_{0}^{r}, B_{0}^{\theta}, B_{0}^{z}\right)(r, z)
\end{array}\right.
$$

First, let us briefly review some results of the MHD system (1.1). For the 2D case, Jiu and Zhao [9] get a global regular solution of MHD with dissipation terms $-(-\triangle)^{\alpha} u$ and $-(-\triangle)^{\beta} B$, when $0 \leq \alpha<\frac{1}{2}, \beta \geq 1,3 \alpha+2 \beta>3$. In particular, they also prove the solution exists globally when $\alpha=0$ and $\beta>\frac{3}{2}$. For more details, one may refer to $[1,6,15,19,20,24]$. In the 3D case, Zhou [23] gets the MHD system with dissipation terms $-(-\triangle)^{\alpha} u$ and $-(-\triangle)^{\beta} B$ satisfying the following conditions:

$$
u(x, t) \in L^{p}\left(0, T ; L^{q}\left(\mathbb{R}^{3}\right)\right), \frac{2 \alpha}{p}+\frac{3}{q} \leq 2 \alpha-1, \frac{3}{2 \alpha-1}<q \leq \infty
$$

or

$$
\Lambda^{\alpha} u(x, t) \in L^{p}\left(0, T ; L^{q}\right), \frac{2 \alpha}{p}+\frac{3}{q} \leq 3 \alpha-1, \frac{3}{3 \alpha-1}<q \leq \frac{3}{\alpha-1}
$$

Then, the solution is globally regular. Other regularity criteria were shown in [4, 7, 8, 21, 22, 25].

When $B=0$, system (1.1) reduce to axisymmetric Navier-Stokes system, there have been extensive studies on the regularity criteria for axisymmetric NavierStokes, cf. $[2,5,10-12,16,18]$. Here, we only mention some results that related to our main results. Firstly, Chae and Lee [2] obtain some regularity criteria for axisymmetric weak solutions of the 3D Navier-Stokes equations with nonzero swirl. In [18], Wei proves the global regularity of solutions to the axially symmetric NavierStokes equations, if $\left\|r u_{\theta}(r, z, 0)\right\|_{L^{\infty}}$ or $\left\|r u_{\theta}(r, z, t)\right\|_{L^{\infty}\left(r \leq r_{0}\right)}$ is smaller than some dimensionless quantity of the initial data. This result improves the one in Lei and Zhang [12]. Motivated by [18], we will study a global regularity for the axially symmetric MHD system.

For equation (1.2), Lei [10] proves the global well-posedness of classical solutions for a family of special axisymmetric initial data whose swirl components of the velocity field and magnetic vorticity field are trivial. In [16], Wang and Wu study the properties of solutions to axially symmetric incompressible MHD system in three dimensions, and construct a family of global smooth solutions by applying the one-dimensional solutions. Recently, Li and Yuan [13] have obtained regularity criteria for the axisymmetric solutions of MHD system, if $\omega_{\theta} \in L^{q}\left(0, T ; L^{p}\left(\mathbb{R}^{3}\right)\right)$, and $n_{\theta} \in L^{q}\left(0, T ; L^{p}\left(\mathbb{R}^{3}\right)\right)$ satisfy

$$
\int_{0}^{T}\left(\left\|\omega_{\theta}\right\|_{p}^{q}+\left\|n_{\theta}\right\|_{p}^{q}\right) d t<\infty, \text { with } \frac{3}{p}+\frac{2}{q} \leq 2, \frac{3}{2}<p \leq \infty, 0<q<\infty
$$


[^0]:    ${ }^{\dagger}$ the corresponding author.
    Email address: xmengliu@zjnu.edu.cn (X. Liu)
    ${ }^{1}$ Department of Mathematics, Zhejiang Normal University, Jinhua, Zhejiang 321004, China

